

LA-UR-19-27161

Approved for public release; distribution is unlimited.

Title: Compatibility Conditions for Quasisymmetry

Author(s): Burby, Joshua William

Intended for: Simons Collaboration webinar

Issued: 2019-07-24

Disclaimer:

Los Alamos National Laboratory, an affirmative action/equal opportunity employer, is operated by Triad National Security, LLC for the National Nuclear Security Administration of U.S. Department of Energy under contract 89233218CNA000001. By approving this article, the publisher recognizes that the U.S. Government retains nonexclusive, royalty-free license to publish or reproduce the published form of this contribution, or to allow others to do so, for U.S. Government purposes. Los Alamos National Laboratory requests that the publisher identify this article as work performed under the auspices of the U.S. Department of Energy. Los Alamos National Laboratory strongly supports academic freedom and a researcher's right to publish; as an institution, however, the Laboratory does not endorse the viewpoint of a publication or guarantee its technical correctness.

Compatibility Conditions for Quasisymmetry

J. W. Burby
(LANL)

July 22nd, 2019
Simons Hour

Supported by LANL LDRD project 20180756PRD4

Quasisymmetry requires equilibrium **and** GC integrability

Definition 1. (quasisymmetric magnetic field)

A magnetic field \mathbf{B} is *quasisymmetric* if

$$(\nabla \times \mathbf{B}) \times \mathbf{B} = \nabla p$$

for some function p **and** the (leading-order) guiding center Lagrangian

$$\begin{aligned} L(\mathbf{X}, v_{\parallel}, \dot{\mathbf{X}}, \dot{v}_{\parallel}) = & (mv_{\parallel} \mathbf{b}(\mathbf{X}) + e\mathbf{A}(\mathbf{X})) \cdot \dot{\mathbf{X}} \\ & - \left(\frac{1}{2}mv_{\parallel}^2 + \mu |\mathbf{B}(\mathbf{X})| \right) \end{aligned}$$

admits a spatial symmetry.

There are various reasons to relax equilibrium constraint.

- Equilibria may support flow

$$\nabla \times (\mathbf{v} \times \mathbf{B}) = 0$$

$$\nabla \cdot (\rho \mathbf{v}) = 0$$

$$\rho \mathbf{v} \cdot \nabla \mathbf{v} + \nabla p = (\nabla \times \mathbf{B}) \times \mathbf{B}$$

There are various reasons to relax equilibrium constraint.

- Equilibria sometimes support **flow**
- Equilibria sometimes support an **anisotropic pressure tensor**

$$\nabla \times (\mathbf{v} \times \mathbf{B}) = 0$$

$$\nabla \cdot (\rho \mathbf{v}) = 0$$

$$\rho \mathbf{v} \cdot \nabla \mathbf{v} + \nabla \cdot \mathbb{P} = (\nabla \times \mathbf{B}) \times \mathbf{B}$$

There are various reasons to relax equilibrium constraint.

- Equilibria sometimes support **flow**
- Equilibria sometimes support an **anisotropic pressure tensor**
- Active **injection of particles or waves** may alter equilibrium force balance

$$\nabla \times (\mathbf{v} \times \mathbf{B}) = 0$$

$$\nabla \cdot (\rho \mathbf{v}) = 0$$

$$\rho \mathbf{v} \cdot \nabla \mathbf{v} + \nabla \cdot \mathbb{P} = (\nabla \times \mathbf{B}) \times \mathbf{B} + \mathbf{F}_{\text{ext}}$$

Weak quasisymmetry requires GC integrability only

Definition 2. (weakly quasisymmetric magnetic field)

A magnetic field \mathbf{B} is *weakly quasisymmetric* if the (leading-order) guiding center Lagrangian

$$L(\mathbf{X}, v_{\parallel}, \dot{\mathbf{X}}, \dot{v}_{\parallel}) = (mv_{\parallel} \mathbf{b}(\mathbf{X}) + e\mathbf{A}(\mathbf{X})) \cdot \dot{\mathbf{X}} \\ - \left(\frac{1}{2}mv_{\parallel}^2 + \mu|\mathbf{B}(\mathbf{X})| \right)$$

admits a spatial symmetry.

We know a little bit about weak quasisymmetry.

weak quasisymmetry: what we know

- ① Quasisymmetry \Rightarrow Weak Quasisymmetry
- ② Weak Quasisymmetry \nRightarrow Quasisymmetry
 - think of non-equilibrium axisymmetric fields

We mostly do not understand weak quasisymmetry.

weak quasisymmetry: open questions

- 1 Are all weakly-quasisymmetric B invariant under rotations and/or translations?

We mostly do not understand weak quasisymmetry.

weak quasisymmetry: open questions

- 1 Are all weakly-quasisymmetric B invariant under rotations and/or translations?
- 2 If not, how to construct non-axisymmetric examples?

We mostly do not understand weak quasisymmetry.

weak quasisymmetry: open questions

- 1 Are all weakly-quasisymmetric \mathbf{B} invariant under rotations and/or translations?
- 2 If not, how to construct non-axisymmetric examples?
- 3 Are quasisymmetric \mathbf{B} perturbations of weakly-quasisymmetric \mathbf{B} ?

We mostly do not understand weak quasisymmetry.

weak quasisymmetry: open questions

- 1 Are all weakly-quasisymmetric \mathbf{B} invariant under **rotations and/or translations**?
- 2 If not, how to construct non-axisymmetric examples?
- 3 Are quasisymmetric \mathbf{B} perturbations of weakly-quasisymmetric \mathbf{B} ?
- 4 Is weak quasisymmetry weak enough to capture eqm flow or anisotropy?

We mostly do not understand weak quasisymmetry.

weak quasisymmetry: open questions

- 1 Are all weakly-quasisymmetric \mathbf{B} invariant under **rotations and/or translations**?
- 2 If not, how to construct non-axisymmetric examples?
- 3 Are quasisymmetric \mathbf{B} perturbations of weakly-quasisymmetric \mathbf{B} ?
- 4 Is weak quasisymmetry weak enough to capture eqm flow or anisotropy?
- 5 Best ways to tackle (1)-(4)?

The purpose of this talk:

Show that weakly-quasisymmetric
fields arise as solutions of a nonlinear
PDE

Why?

- Provide a concrete description of weakly-quasisymmetric fields
 - the definition is rather abstract
- Introduce familiar framework to begin addressing open questions about weak quasisymmetry

The PDE will be derived as follows

Step One:

Identify PDE for generator \mathbf{u} of weak quasisymmetry

Step Two:

Extract PDE for \mathbf{B} as compatibility conditions for \mathbf{u}
equation to have solution

First step is straightforward Lagrangian mechanics

Theorem 1. (conditions for weak quasisymmetry)

A magnetic field \mathbf{B} is weakly-quasisymmetric if and only if there is a vector field \mathbf{u} such that

$$(\nabla \times \mathbf{B}) \times \mathbf{u} + \nabla(\mathbf{u} \cdot \mathbf{B}) = 0$$

$$\nabla \times (\mathbf{B} \times \mathbf{u}) = 0$$

$$\nabla \cdot \mathbf{u} = 0.$$

See:

- Burby, Qin, “Toroidal precession as a geometric phase,” *Phys. Plasmas* **20**, 012511 (2013).
- R. S. MacKay’s talk from first Simon’s Meeting

Basic idea:

Find conditions on B ensuring solution for u exists

Basic idea:

Find conditions on B ensuring solution for u exists

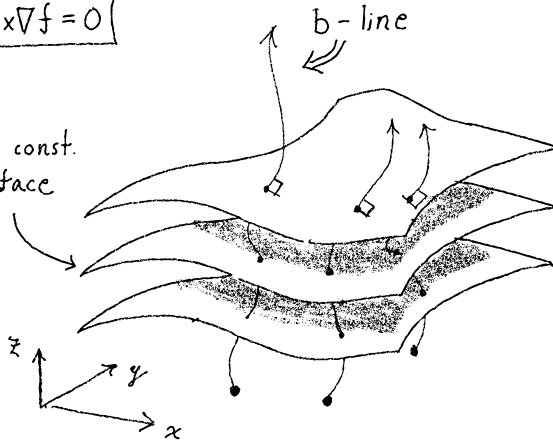
How do we find such compatibility conditions?

Illustrative example

Surfaces perpendicular to a magnetic field

$$\mathbf{b} \times \nabla f = 0$$

$f = \text{const.}$
surface



Perp. surfaces satisfy overdetermined PDE.

$$\mathbf{b}(\mathbf{x}) \times \nabla f(\mathbf{x}) = 0, \quad \forall \mathbf{x}, \nabla f(\mathbf{x}) \neq 0$$

Three equations; one unknown function f !

\Rightarrow We should expect compatibility conditions for a solution to exist

Equality of mixed partials gives compatibility condition.

$$0 = \nabla \cdot (\mathbf{b} \times \nabla f) = (\nabla \times \mathbf{b}) \cdot \nabla f$$

\Downarrow (because a solution must satisfy $\nabla f = \lambda \mathbf{b}$) \Downarrow

$$\tau = \mathbf{b} \cdot \nabla \times \mathbf{b} = 0$$

Equality of mixed partials gives compatibility condition.

$$0 = \nabla \cdot (\mathbf{b} \times \nabla f) = (\nabla \times \mathbf{b}) \cdot \nabla f$$

\Downarrow (because a solution must satisfy $\nabla f = \lambda \mathbf{b}$) \Downarrow

$$\tau = \mathbf{b} \cdot \nabla \times \mathbf{b} = 0$$

This is a compatibility condition!

Frobenius Thm says this is only compatibility condition.

Theorem (Frobenius)

- a If a solution f exists, $\tau = 0$ everywhere.
- b Conversely, if $\tau = 0$ everywhere, a smooth solution exists in a neighborhood of each \mathbf{x} .

Q: How can we generalize this example to treat weak quasisymmetry?

A: Resort to exterior differential systems (EDS)
theory!

EDS theory systematizes finding **all*** compatibility conditions for **any** PDE.

CC extraction procedure:

(0) Reformulate PDE as vanishing of differential forms θ_i

$$\mathbf{b} \times \nabla f = 0$$

$$\Updownarrow$$

$$\theta = (\mathbf{b} \cdot d\mathbf{x}) \wedge df = 0$$

EDS theory systematizes finding **all*** compatibility conditions for **any** PDE.

CC extraction procedure:

- (0) Reformulate PDE as vanishing of differential forms θ_i
- (1) For each dependent variable Q , set $dQ = \mathbf{p}^{(Q)} \cdot d\mathbf{x}$ in θ_i , $d\theta_i$; find constraints on $\mathbf{p}^{(Q)}$'s. (\mathbf{x} = independent variables)

$$\begin{aligned}\theta &= (\mathbf{b} \cdot d\mathbf{x}) \wedge (\mathbf{p}^{(f)} \cdot d\mathbf{x}) = 0 \\ d\theta &= (\nabla \times \mathbf{b} \cdot d\mathbf{S}) \wedge (\mathbf{p}^{(f)} \cdot d\mathbf{x}) = 0 \\ &\quad \Updownarrow \\ \mathbf{b} \times \mathbf{p}^{(f)} &= 0 \\ (\nabla \times \mathbf{b}) \cdot \mathbf{p}^{(f)} &= 0\end{aligned}$$

The \mathbf{p} 's may be interpreted as derivatives

EDS theory systematizes finding **all*** compatibility conditions for **any** PDE.

CC extraction procedure:

- (0) Reformulate PDE as vanishing of differential forms θ_i
- (1) For each dependent variable Q , set $dQ = \mathbf{p}^{(Q)} \cdot d\mathbf{x}$ in $\theta_i, d\theta_i$; find constraints on $\mathbf{p}^{(Q)}$'s. (\mathbf{x} = independent variables)
- (2) If \mathbf{p} -equations constrain dependent variables, sub constraint into θ_i then return to (1). Otherwise move on.

f does not appear, so nothing to do!

EDS theory systematizes finding **all*** compatibility conditions for **any** PDE.

CC extraction procedure:

- (0) Reformulate PDE as vanishing of differential forms θ_i
- (1) For each dependent variable Q , set $dQ = \mathbf{p}^{(Q)} \cdot d\mathbf{x}$ in θ_i , $d\theta_i$; find constraints on $\mathbf{p}^{(Q)}$'s. (\mathbf{x} = independent variables)
- (2) If \mathbf{p} -equations constrain dependent variables, sub constraint into θ_i then return to (1). Otherwise move on.
- (3) If \mathbf{p} -equations constrain independent variables, impose as CC

$$\mathbf{b} \times \mathbf{p}^{(f)} = 0, \quad (\nabla \times \mathbf{b}) \cdot \mathbf{p}^{(f)} = 0$$

$$\Updownarrow$$

$$\mathbf{p}^{(f)} = \lambda \mathbf{b}, \quad (\nabla \times \mathbf{b}) \cdot \mathbf{p}^{(f)} = 0$$

$$\Updownarrow$$

$$\mathbf{p}^{(f)} = \lambda \mathbf{b}, \quad \mathbf{b} \cdot (\nabla \times \mathbf{b})(\mathbf{x}) = 0$$

EDS theory systematizes finding **all*** compatibility conditions for **any** PDE.

CC extraction procedure:

- (0) Reformulate PDE as vanishing of differential forms θ_i
- (1) For each dependent variable Q , set $dQ = \mathbf{p}^{(Q)} \cdot d\mathbf{x}$ in $\theta_i, d\theta_i$; find constraints on $\mathbf{p}^{(Q)}$'s. (\mathbf{x} = independent variables)
- (2) If \mathbf{p} -equations constrain dependent variables, sub constraint into θ_i then return to (1). Otherwise move on.
- (3) If \mathbf{p} -equations constrain independent variables, impose as CC
- (4) If \mathbf{p} -equations solvable for each \mathbf{x} , apply *prolongation*. Otherwise return to (2).

$$\theta = (\mathbf{b} \cdot d\mathbf{x}) \wedge df \quad \rightarrow \quad \Theta = df - \lambda \mathbf{b} \cdot d\mathbf{x}$$

Prolongation differentiates the PDE in an intelligent way

EDS theory systematizes finding **all*** compatibility conditions for **any** PDE.

CC extraction procedure:

- (0) Reformulate PDE as vanishing of differential forms θ_i
- (1) For each dependent variable Q , set $dQ = \mathbf{p}^{(Q)} \cdot d\mathbf{x}$ in $\theta_i, d\theta_i$; find constraints on $\mathbf{p}^{(Q)}$'s. (\mathbf{x} = independent variables)
- (2) If \mathbf{p} -equations constrain dependent variables, sub constraint into θ_i then return to (1). Otherwise move on.
- (3) If \mathbf{p} -equations constrain independent variables, impose as CC
- (4) If \mathbf{p} -equations solvable for each \mathbf{x} , apply *prolongation*. Otherwise return to (2).
- (5) Apply “Cartan’s test.” If pass, stop. If fail, return to (1).

Cartan’s test amounts to linear algebra. **Will not discuss here.**

EDS theory systematizes finding **all*** compatibility conditions for **any** PDE.

*Important Technicalities

- CCs necessary conditions for *any* solution to exist
- All CCs satisfied \Rightarrow formal power series solutions exist
- If PDE coefficients real analytic, formal power series converge in small domain. (Generalized Cauchy-Kowalevski Thm.)
- Sometimes you're lucky! Satisfying all CCs may give exactly-solvable system.
 - This happens for weak quasisymmetry!

Preceding procedure produces all compatibility conditions for \mathbf{u} -equation.

Theorem 2. (compatibility conditions for weak quasisymmetry)

A non-vacuum magnetic field \mathbf{B} is weakly-quasisymmetric if and only if there are potentials φ, ψ such that

$$\nabla\varphi = (\nabla \times \mathbf{B}) \times \mathbf{e}$$

$$\nabla\psi = \mathbf{B} \times e^{-\varphi} \mathbf{e}$$

$$0 = \mathbf{e} \cdot \nabla(\mathbf{B} \cdot \nabla|\mathbf{B}| \times \nabla\tau)$$

where $\tau = \mathbf{b} \cdot \nabla \times \mathbf{b}$ and the vector field \mathbf{e} is given by

$$\mathbf{e} = \frac{\nabla|\mathbf{B}| \times \nabla\tau}{\mathbf{B} \cdot \nabla|\mathbf{B}| \times \nabla\tau}.$$

Preceding procedure produces all compatibility conditions for \mathbf{u} -equation.

Theorem 2. (compatibility conditions for weak quasisymmetry)

A non-vacuum magnetic field \mathbf{B} is weakly-quasisymmetric if and only if there are potentials φ, ψ such that

$$\nabla\varphi = (\nabla \times \mathbf{B}) \times \mathbf{e}$$

$$\nabla\psi = \mathbf{B} \times e^{-\varphi} \mathbf{e}$$

$$0 = \mathbf{e} \cdot \nabla(\mathbf{B} \cdot \nabla|\mathbf{B}| \times \nabla\tau)$$

where $\tau = \mathbf{b} \cdot \nabla \times \mathbf{b}$ and the vector field \mathbf{e} is given by

$$\mathbf{e} = \frac{\nabla|\mathbf{B}| \times \nabla\tau}{\mathbf{B} \cdot \nabla|\mathbf{B}| \times \nabla\tau}.$$

Any 3D solution of this PDE will be non-axisymmetric weakly quasisymmetric field.

1. Near-axis expansion of PDE for weak-quasisymmetry

10246v2 [physics.plasm-ph] 4 Dec 2018

Under consideration for publication in J. Plasma Phys.

1

Direct construction of optimized stellarator shapes. II. Numerical quasisymmetric solutions

Matt Landreman¹†, Wrick Sengupta² and Gabriel G Plunk³

¹Institute for Research in Electronics and Applied Physics, University of Maryland, College Park MD 20742, USA

²Courant Institute of Mathematical Sciences, New York University, New York NY 10012, USA

³Max Planck Institute for Plasma Physics, Greifswald, Germany

(Received xx; revised xx; accepted xx)

Quasisymmetric stellarators are appealing intellectually and as fusion reactor candidates since the guiding center particle trajectories and neoclassical transport are isomorphic to those in a tokamak, implying good confinement. Previously, quasisymmetric magnetic fields have been identified by applying black-box optimization algorithms to minimize symmetry-breaking Fourier modes of the field strength B . Here instead we directly construct magnetic fields in cylindrical coordinates that are quasisymmetric to leading order in distance from the magnetic axis, without using optimization. The method involves solution of a 1-dimensional nonlinear ordinary differential equation, originally derived by Garren and Boozer [*Phys. Fluids B* **3**, 2805 (1991)]. We demonstrate the usefulness and accuracy of this optimization-free approach by providing the results of this construction as input to the codes VMEC and BOOZ_XFORM, confirming the purity and scaling of the magnetic spectrum. The space of magnetic fields that are quasisymmetric to this order is parameterized by the magnetic axis shape along with three other real numbers, one of which reflects the on-axis toroidal current density, and another one of which is zero for stellarator symmetry. The method here could be used to generate good initial conditions for conventional optimization, and its speed enables exhaustive searches of parameter space.

Where do we go from here?

2. Even weaker quasisymmetry

When $\epsilon = \rho/L = 0$, the GC Lagrangian blows up...

$$L(\mathbf{X}, v_{\parallel}, \dot{\mathbf{X}}, \dot{v}_{\parallel}) = (mv_{\parallel} \mathbf{b}(\mathbf{X}) + \frac{1}{\epsilon} e \mathbf{A}(\mathbf{X})) \cdot \dot{\mathbf{X}} \\ - \left(\frac{1}{2} m v_{\parallel}^2 + \mu |\mathbf{B}(\mathbf{X})| \right)$$

Where do we go from here?

2. Even weaker quasisymmetry

...but the GC Poisson bracket and Hamiltonian do not.

$$\{f, g\} = (\mathbf{b} \cdot \nabla f) \partial_{v_{\parallel}} g - \partial_{v_{\parallel}} f (\mathbf{b} \cdot \nabla g) + O(\epsilon)$$

$$H = \frac{1}{2} m v_{\parallel}^2 + \mu |\mathbf{B}| + O(\epsilon)$$

Where do we go from here?

2. Even weaker quasisymmetry

Theorem 3 (Conditions for symmetry of leading-order Hamiltonian GC dynamics)

The leading-order GC Poisson bracket and Hamiltonian admit a spatial symmetry if and only if there is a \mathbf{u} such that

$$\begin{aligned}\mathbf{u} \cdot \nabla |\mathbf{B}| &= 0 \\ \mathbf{u} \cdot \nabla \mathbf{B} - \mathbf{B} \cdot \nabla \mathbf{u} &= 0.\end{aligned}$$

(N. B. These conditions are satisfied automatically assuming weak quasisymmetry.)

What are the compatibility conditions on \mathbf{B} to ensure \mathbf{u} exists?

Where do we go from here?

3. Continue to improve understanding of weak quasisymmetry

- Switch roles of \mathbf{B} and \mathbf{u} to find compatibility conditions on \mathbf{u} .
 - e.g. if $\nabla \mathbf{u} + (\nabla \mathbf{u})^T = 0$ is compatibility condition, then *all weakly quasisymmetric fields must be axisymmetric.*

Where do we go from here?

3. Continue to improve understanding of weak quasisymmetry

- Switch roles of \mathbf{B} and \mathbf{u} to find compatibility conditions on \mathbf{u} .
 - e.g. if $\nabla \mathbf{u} + (\nabla \mathbf{u})^T = 0$ is compatibility condition, then *all weakly quasisymmetric fields must be axisymmetric.*
- If weak quasisymmetric fields exist, can they support flow or anisotropic pressure?

Where do we go from here?

3. Continue to improve understanding of weak quasisymmetry

- Switch roles of \mathbf{B} and \mathbf{u} to find compatibility conditions on \mathbf{u} .
 - e.g. if $\nabla \mathbf{u} + (\nabla \mathbf{u})^T = 0$ is compatibility condition, then *all weakly quasisymmetric fields must be axisymmetric.*
- If 3D weak quasisymmetric fields **possible**, can it support flow or anisotropic pressure?
- If 3D weak quasisymmetry **impossible**, how nearly can it be satisfied?

END